Aspects of Non-Supersymmetric String Phenomenology

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I. R and I. Florakis, arXiv:1608.04582 [hep-th], Nucl. Phys. B 913 (2016) 495 arXiv:1703.09272 [hep-th], Nucl. Phys. B 921 (2017) 1

Outline

- Introduction
- Non supersymmetric string models Coordinate dependent compactifications
- · A class of semi-realistic non supersymmetric string vacua
- One loop potential for the moduli fields Cosmological constant
- Gauge coupling thresholds The decompactification problem
- Conclusions

The Standard Model

The Standard Model of particle interactions is a very successful theory.

However, it leaves a number of unanswered questions (Mass origin, flavor puzzle, charge quantization, a number of parameters, dark matter, hierarchy problem, gravity...)

Supersymmetry has been introduced to provide a solution to the gauge hierarchy problem and guarantee stability towards quantum corrections without fine-tuning. The introduction of SUSY at a few TeV leads also to coupling unification.

If SUSY were an exact symmetry of the nature every particle and its superpartener would have degenerate masses. However, this is not verified experimentally so SUSY must be broken.

Non-supersymmetric strings

Space-time supersymmetry is not required for consistency in string theory.

From the early days of the first string revolution it was known that heterotic strings comprise the SUSY $E_8 \times E_8$ and SO(32) models as well as the non-supersymmetric tachyon free $SO(16) \times SO(16)$ theory.

However, non-supersymmetric string phenomenology has not received much attention until recently.

see e.g.

S. Abel, K. R. Dienes and E. Mavroudi (2015,2017), J. R. and I. Florakis (2016,2017), Y. Sugawara, T. Wada (2016), A. Lukas, Z. Lalak and E. E. Svanes (2015), S.G. Nibbelink, O. Loukas, A. Mütter, E. Parr, P. K. S. Vaudrevange (2017)

SUSY breaking in String Theory

Any scenario of supersymmetry breaking in the context of string theory has to address some important issues, as

- Resolve M_W/M_P hierarchy
- Compatibility with gauge coupling evolution (unification)
- Account for the smallness of the cosmological constant
- Resolve possible instabilities (tachyons)
- Moduli field stabilisation

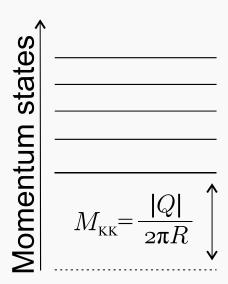
Coordinate dependent compactifications

A stringy Scherk–Schwartz mechanism involves an extra dimension X^5 and a conserved charge Q. Upon compactification

$$\Phi\left(X^{5}+2\pi R\right)=e^{iQ}\Phi\left(X^{5}\right)$$

we obtain a shifted tower of Kaluza–Klein states for charged fields, starting at $M_{KK} = \frac{|Q|}{2\pi R}$

$$\Phi(X^5) = e^{\frac{i QX^5}{2\pi R}} \sum \Phi_n e^{i nX^5/R}$$



see e.g. J. Scherk and J. H. Schwarz (1978,1979), R. Rohm (1984), C. Kounnas and M. Porrati (1988), S. Ferrara, C. Kounnas, M. Porrati and F. Zwirner (1989), C. Kounnas and B. Rostand, (1990)

Coordinate dependent compactifications

 $Q = \text{Fermion number} \Rightarrow \text{leads to different masses for}$ fermions-bosons (lying in the same supermultiplet) and thus to spontaneous breaking of supersymmetry.

SUSY breaking related to the compactification radius $M \sim \frac{1}{R}$

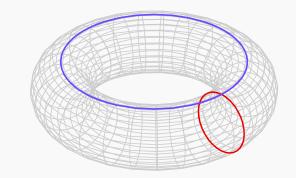
We are going to consider the implementation of this mechanism in a class of phenomenologically interesting *SO*(10) heterotic string models.

Gravitino mass

We consider compactifications of the six internal dimensions in three separate two-tori parametrised by the $T^{(i)}, U^{(i)}, i = 1, 2, 3$ moduli. For simplicity, we will consider realising the Scherk–Schwartz mechanism utilising the $T^{(1)}, U^{(1)}$ torus.

At tree level the gravitino receives a mass

$$m_{3/2} = \frac{|U^{(1)}|}{\sqrt{T_2^{(1)}U_2^{(1)}}} = \frac{1}{R_1}$$



for a square torus: $T = iR_1 R_2$, $U = iR_2/R_1$ All $T^{(i)}$, $U^{(i)}$ moduli remain massless.

At $R_1 \to \infty$ we have $m_{3/2} = 0$ and the supersymmetry is restored.

One loop potential

The effective potential at one loop, as a function moduli $t_l = T^{(i)}, U^{(i)}$, is obtained by integrating the string partition function $Z(\tau_1, \tau_2; t_l)$ over the worldsheet torus Σ_1

$$V_{\mathrm{one-loop}}(t_I) = -\frac{1}{2(2\pi)^4} \int_{\mathcal{F}} \frac{d^2\tau}{\tau_2^3} Z(\tau, \bar{\tau}; t_I),$$

 \mathcal{T}_{2} \mathcal{F} -1/2 +1/2 \mathcal{T}_{1}

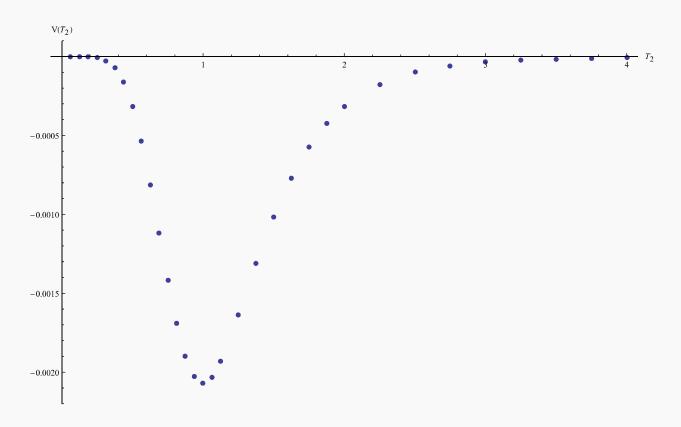
where $\tau = \tau_1 + i\tau_2$ and \mathcal{F} is the fundamental domain .

For given values of the moduli

$$Z = \sum_{\substack{n \in \mathbb{Z}/2 \\ n \ge -1/2}} \sum_{m \in \mathbb{Z}} Z_{n,m} \, q_{\mathbf{r}}^n \, q_{\mathbf{i}}^m = \sum_{\substack{n \in \mathbb{Z}/2 \\ n \ge -1/2}} \left[\sum_{m=-[n]-1}^{[n]+2} Z_{n,m} \, q_{\mathbf{i}}^m \right] \, q_{\mathbf{r}}^n \, .$$

where $q_{\rm r}=e^{-2\pi\tau_2}$ and $q_{\rm i}=e^{2\pi i \tau_1}$

One loop moduli potentials



Typical one-loop potential as a function of the modulus $T_2 = R^2$.

Undesirable features: SUSY breaking at the string scale, huge cosmological constant, region of tachyon instabilities

One loop potential: Analytic results

$$Z = \frac{1}{2^{8}} \frac{1}{\eta^{12} \bar{\eta}^{24}} \sum_{H_{1}, G_{1} = 0, 1} \Gamma_{2,2}^{\text{shift}} \begin{bmatrix} H_{1} \\ G_{1} \end{bmatrix} \left(T^{(1)}, U^{(1)} \right)$$

$$\times \sum_{h_{2}, H = 0, 1 \atop g_{2}, G = 0, 1} \sum_{\ell, \sigma, \delta_{3}, \delta_{4} = 0, 1} (-1)^{\hat{\Phi}'} \times \vartheta \begin{bmatrix} 1 + H_{1} + h_{2} \\ 1 + G_{1} + g_{2} \end{bmatrix}^{2} \vartheta \begin{bmatrix} 1 + H_{1} \\ 1 + G_{1} \end{bmatrix}^{2}$$

$$\times \bar{\vartheta} \begin{bmatrix} k \\ \ell \end{bmatrix}^{6} \bar{\vartheta} \begin{bmatrix} k + h_{2} \\ \ell + g_{2} \end{bmatrix}^{2} \bar{\vartheta} \begin{bmatrix} \rho \\ \sigma \end{bmatrix}^{4} \bar{\vartheta} \begin{bmatrix} \rho + H \\ \sigma + G \end{bmatrix}^{4} \vartheta \begin{bmatrix} \gamma_{2} \\ \delta_{2} \end{bmatrix} \vartheta \begin{bmatrix} \gamma_{2} + h_{2} \\ \delta_{2} + g_{2} \end{bmatrix}$$

$$\times \bar{\vartheta} \begin{bmatrix} \gamma_{2} \\ \delta_{2} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} \gamma_{2} + h_{2} \\ \delta_{2} + g_{2} \end{bmatrix} \vartheta \begin{bmatrix} \gamma_{3} - h_{3} \\ \delta_{3} - g_{3} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} \gamma_{3} - h_{3} \\ \delta_{3} - g_{3} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} \gamma_{3} - h_{3} \\ \delta_{3} - g_{3} \end{bmatrix}$$

One loop potential: Asymptotic limit $R \to \infty$

The asymptotic behaviour of the potential is

$$V_{\text{one-loop}}(R) = -\frac{(n_B - n_F)}{2^4 \pi^7 R^4} \sum_{m_1, m_2 \in \mathbb{Z}} \frac{U_2^3}{\left| m_1 + \frac{1}{2} + U \, m_2 \right|^6} + \mathcal{O}\left(e^{-\sqrt{2\pi}R}\right)$$

$$V_{\text{one-loop}}(R) = \xi \frac{(n_B - n_F)}{R^4} + \text{exponentially supressed}$$

where ξ is a constant and n_B , n_F stand for the number of bosonic and fermionic degrees of freedom respectively.

Super no scale models $n_B = n_F$. Cosmological constant is exponentially small.

C. Kounnas and H. Partouche (2016), T. Coudarchet and H. Partouche (2018)

A class of models

Consider a big class of semi-realistic $Z_2 \times Z_2$ heterotic string vacua for explicit realisations of the Scherk–Schwarz scenario. Study chirality, moduli potential and thresholds.

To this end we utilise both the free fermionic formulation and orbifold formulation. In the former we have full control of the spectrum in the latter we have explicit moduli dependence.

In the free fermionic formulation we use the model classification techniques developed in

A. Gregori, C. Kounnas and J. R. (1999)

A. E. Faraggi, C. Kounnas, S. E. M. Nooij and J. R. (2004)

A. E. Faraggi, C. Kounnas and J. R. (2007)

The class of SO(10) models

We consider a class of four dimensional N=1 heterotic models spontaneously broken to N=0 via the Scherk–Schwarz mechanism. At technical level, this class is generated by 9 basis vectors in the free fermionic formulation and is parametrised by a set of 36 phases associated with generalised GSO projections. It comprises $2^{9(9-1)/2+1} \sim 10^{11}$ (a priori) distinct models.

The $E_8 \times E_8$ gauge symmetry is reduced to

$$SO(10) \times SO(8)^2 \times U(1)^2$$

We select models using the following criteria

- absence of tachyons
- SO(10) chirality
- compatibility with Scherk–Schwarz breaking of N = 1 SUSY

Class of models: Basis vectors

The free fermions in the light-cone gauge are:

left:
$$\psi^{\mu}, \chi^{1,...,6}, \quad y^{1,...,6}, \omega^{1,...,6}$$

right: $\bar{y}^{1,...,6}, \bar{\omega}^{1,...,6}, \quad \bar{\eta}^{1,2,3}, \bar{\psi}^{1,...,5}, \quad \bar{\phi}^{1,...,8}$
The class of vacua under consideration is defined by $\beta_1 = 1 = \{\psi^{\mu}, \chi^{1,...,6}, y^{1,...,6}, \omega^{1,...,6}| \bar{y}^{1,...,6}, \bar{\omega}^{1,...,6}, \bar{\eta}^{1,2,3}, \bar{\psi}^{1,...,5}, \bar{\phi}^{1,...,8} \}$
 $\beta_2 = S = \{\psi^{\mu}, \chi^{1,...,6}\}$
 $\beta_3 = T_1 = \{y^{12}, \omega^{12}| \bar{y}^{12}, \bar{\omega}^{12} \}$
 $\beta_4 = T_2 = \{y^{34}, \omega^{34}| \bar{y}^{34}, \bar{\omega}^{34} \}$
 $\beta_5 = T_3 = \{y^{56}, \omega^{56}| \bar{y}^{56}, \bar{\omega}^{56} \}$
 $\beta_6 = b_1 = \{\chi^{34}, \chi^{56}, y^{34}, y^{56}| \bar{y}^{34}, \bar{y}^{56}, \bar{\psi}^{1,...,5}, \bar{\eta}^1 \}$
 $\beta_7 = b_2 = \{\chi^{12}, \chi^{56}, y^{12}, y^{56}| \bar{y}^{12}, \bar{y}^{56}, \bar{\psi}^{1,...,5}, \bar{\eta}^2 \}$
 $\beta_8 = z_1 = \{\bar{\phi}^{1,...,4}\}$
 $\beta_9 = z_2 = \{\bar{\phi}^{5,...,8}\}$
and a variable set of $9(9-1)/2+1=37$ phases $c[\frac{\beta_i}{\beta_i}]$.

Chirality

Fermion generations, transforming as SO(10) spinorials, arise from $B_{pq}^{I} = S + b_{pq}^{I}$, I = 1, 2, 3 where $b_{pq}^{1} = b^{1} + p T_{2} + q T_{3}$, $b_{pq}^{2} = b^{2} + p T_{1} + q T_{2}$, $b_{pq}^{3} = x + b^{1} + b^{2} + p T_{1} + q T_{2}$, with $p, q \in \{0, 1\}$, and $x = 1 + S + \sum_{i=1}^{3} T_{i} + \sum_{k=1}^{2} z_{k}$.

Number of generations $N = \sum_{l=1,2,3} \chi^l$ where

$$\chi_{pq}^{1} = -4c \begin{bmatrix} B_{pq}^{1} \\ S + b_{2} + (1 - q)T_{3} \end{bmatrix} P_{pq}^{1},$$

$$\chi_{pq}^{2} = -4c \begin{bmatrix} B_{pq}^{2} \\ S + b_{1} + (1 - q)T_{3} \end{bmatrix} P_{pq}^{2},$$

$$\chi_{pq}^{3} = -4c \begin{bmatrix} B_{pq}^{3} \\ S + b_{1} + (1 - q)T_{1} \end{bmatrix} P_{pq}^{3},$$

and

$$P_{pq}^{l} = \frac{1}{2^{3}} \left(1 - c \begin{bmatrix} B_{pq}^{l} \\ T_{l} \end{bmatrix} \right) \left(1 - c \begin{bmatrix} B_{pq}^{l} \\ Z_{1} \end{bmatrix} \right) \left(1 - c \begin{bmatrix} B_{pq}^{l} \\ Z_{2} \end{bmatrix} \right)$$

Orbifold Partition function

The one-loop partition function at the generic point reads

$$Z = \frac{1}{\eta^{12}\bar{\eta}^{24}} \frac{1}{2^{3}} \sum_{\substack{h_{1},h_{2},H\\g_{1},g_{2},G}} \frac{1}{2^{3}} \sum_{\substack{a,k,\rho\\b,\ell,\sigma}} \frac{1}{2^{3}} \sum_{\substack{H_{1},H_{2},H_{3}\\G_{1},G_{2},G_{3}}} (-1)^{a+b+HG+\Phi}$$

$$\times \vartheta\begin{bmatrix} a\\b\end{bmatrix} \vartheta\begin{bmatrix} a+h_{1}\\b+g_{1} \end{bmatrix} \vartheta\begin{bmatrix} a+h_{2}\\b+g_{2} \end{bmatrix} \vartheta\begin{bmatrix} a-h_{1}-h_{2}\\b-g_{1}-g_{2} \end{bmatrix}$$

$$\times \Gamma^{(1)}_{2,2} \begin{bmatrix} H_{1}\\G_{1} \end{bmatrix} \begin{bmatrix} h_{1}\\g_{1} \end{bmatrix} (T^{(1)}, U^{(1)}) \Gamma^{(2)}_{2,2} \begin{bmatrix} H_{2}\\G_{2} \end{bmatrix} \begin{bmatrix} h_{2}\\g_{2} \end{bmatrix} (T^{(2)}, U^{(2)}) \Gamma^{(3)}_{2,2} \begin{bmatrix} H_{3}\\G_{3} \end{bmatrix} \begin{bmatrix} h_{1}+h_{2}\\g_{1}+g_{2} \end{bmatrix} (T^{(3)}, U^{(3)})$$

$$\times \bar{\vartheta} \begin{bmatrix} k\\\ell+g_{1} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} k+h_{1}\\\ell+g_{1} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} k+h_{2}\\\ell+g_{2} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} k-h_{1}-h_{2}\\\ell-g_{1}-g_{2} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} \rho\\\ell-g_{1}-g_{2} \end{bmatrix} \bar{\vartheta} \begin{bmatrix} \rho+H\\g_{1}-g_{2} \end{bmatrix}^{4}$$

Where $T^{(i)} = T_1^{(i)} + iT_2^{(i)}$, $U^{(i)} = U_1^{(i)} + iU_2^{(i)}$ are the moduli of the three two tori, $\eta(\tau)$ is the Dedekind eta function and $\vartheta[^{\alpha}_{\beta}](\tau)$ stand for the Jacobi theta functions.

Connection with fermionic formulation

Fermionic point
$$T = i$$
 and $U = (1 + i)/2$
Phase $\Phi\left(c\begin{bmatrix}\beta_i\\\beta_j\end{bmatrix}\right)$

Twisted/shifted lattices

$$\Gamma_{2,2}[^{H_i}_{G_i}|^h_g](T,U) = \begin{cases} \left|\frac{2\eta^3}{\vartheta[^{1-h}_{1-g}]}\right|^2 &, (H_i,G_i) = (0,0) \text{ or } (H_i,G_i) = (h,g) \\ \Gamma^{\text{shift}}_{2,2}[^{H_i}_{G_i}](T,U) &, h = g = 0 \\ 0 &, \text{ otherwise} \end{cases}$$

$$\Gamma_{2,2}^{\text{shift}}[_{G_i}^{H_i}](T,U) = \sum_{\substack{m_1,m_2\\n_1,n_2}} (-1)^{G(m_1+n_2)} q^{\frac{1}{4}|P_L|^2} \bar{q}^{\frac{1}{4}|P_R|^2},$$

with

$$P_{L} = \frac{m_{2} + \frac{H_{i}}{2} - Um_{1} + T(n_{1} + \frac{H_{i}}{2} + Un_{2})}{\sqrt{T_{2}U_{2}}},$$

$$P_{R} = \frac{m_{2} + \frac{H_{i}}{2} - Um_{1} + \overline{T}(n_{1} + \frac{H_{i}}{2} + Un_{2})}{\sqrt{T_{2}U_{2}}}.$$

Typical partition functions

Some typical expansions of partition functions (fermionic point)

$$\begin{split} Z_{(A)} &= \frac{2q_{\rm i}}{q_{\rm r}} - \frac{16q_{\rm i}}{\sqrt{q_{\rm r}}} + \left(-312 + 32q_{\rm i} + 56q_{\rm i}^2\right) \\ &\quad + \left(4064 + \frac{6144}{q_{\rm i}} + 512q_{\rm i} - 416q_{\rm i}^2\right) \sqrt{q_{\rm r}} \\ &\quad + \left(12288 + \frac{16384}{q_{\rm i}^2} + \frac{103680}{q_{\rm i}} - 12320q_{\rm i} - 256q_{\rm i}^2 + 792q_{\rm i}^3\right) q_{\rm r} + \dots \\ Z_{(B)} &= \frac{2q_{\rm i}}{q_{\rm r}} - \frac{32q_{\rm i}}{\sqrt{q_{\rm r}}} + \left(8 + 224q_{\rm i} + 56q_{\rm i}^2\right) + \left(1984 + \frac{2048}{q_{\rm i}} - 1024q_{\rm i} - 832q_{\rm i}^2\right) \sqrt{q_{\rm r}} \\ &\quad + \left(30720 + \frac{10240}{q_{\rm i}^2} + \frac{92160}{q_{\rm i}} + 1760q_{\rm i} + 5376q_{\rm i}^2 + 792q_{\rm i}^3\right) q_{\rm r} + \dots \\ Z_{(C)} &= \frac{2q_{\rm i}}{q_{\rm r}} - \frac{16q_{\rm i}}{\sqrt{q_{\rm r}}} + \left(40 + 64q_{\rm i} + 56q_{\rm i}^2\right) + \left(224 + \frac{6912}{q_{\rm i}} + 768q_{\rm i} - 672q_{\rm i}^2\right) \sqrt{q_{\rm r}} \\ &\quad + \left(14336 + \frac{9216}{q_{\rm i}^2} + \frac{118656}{q_{\rm i}} - 10144q_{\rm i} + 3072q_{\rm i}^2 + 792q_{\rm i}^3\right) q_{\rm r} + \dots \end{split}$$

Classification

A comprehensive computer scan results in 7×10^4 models that satisfy all criteria.

We expand the partition function in powers of $q_{\rm r}=e^{-2\pi\tau_2}$

$$Z = \sum_{\substack{n \in \mathbb{Z}/2 \\ n > -1/2}} W_n \, q_{\mathbf{r}}^n$$

The constant term at the fermionic point W_0 or the generic point W_0^G is proportional to $n_B - n_F$.

	$W_0 < 0$	$W_0 = 0$	$W_0 > 0$
$W_0^G < 0$	3560	0	1856
$W_0^G=0$	96	0	8848
$W_0^G > 0$	0	0	62192
Total	3656	0	72896

Table 1: Number of chiral models for the subclasses of models with

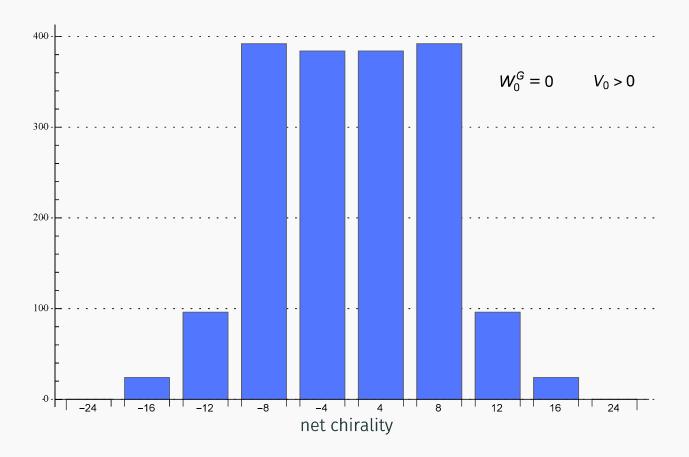
Numerical calculation

n	Model A	Model B
1	24.4	24.4
$-\frac{1}{2}$	-9.87	-19.7
0	172.	2.11
<u>1</u>	-29.6	-17.7
1	3.13	-2.73
<u>3</u>	9.71	8.18
Total	+170.	-5.47

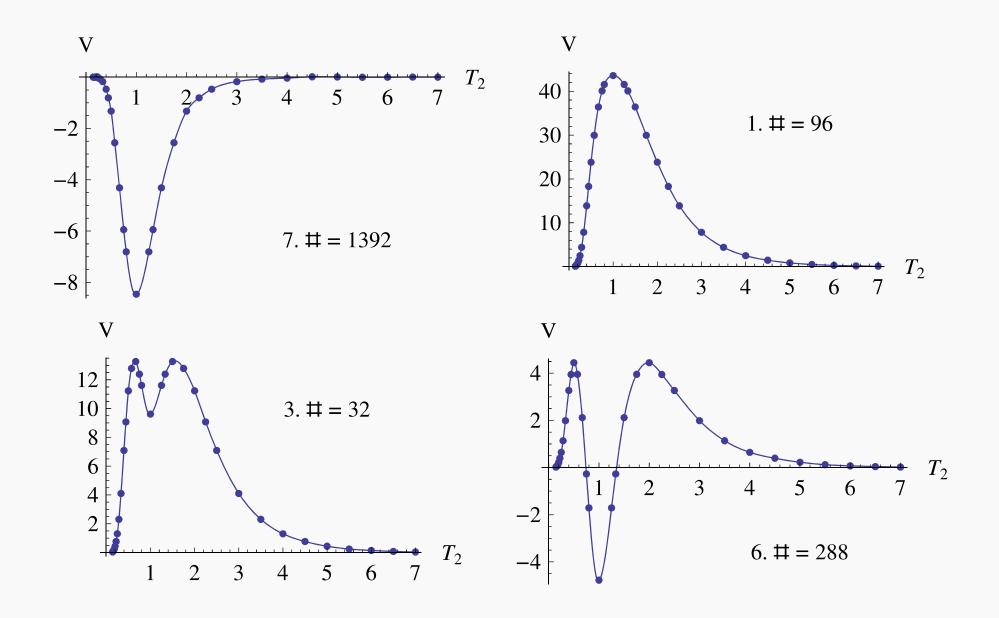
Contributions to the one-loop potential $2(2\pi)^4V_{1-loop}$ arranged according to energy level for two models (A and B) at the fermionic point.

Chiral super no-scale models

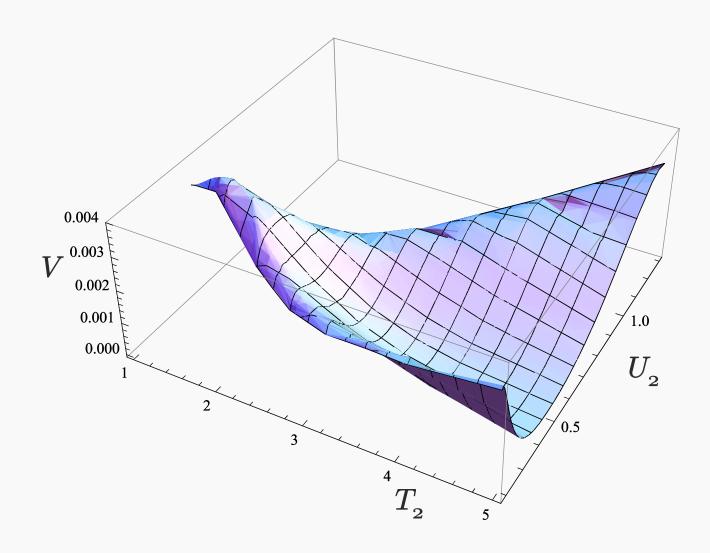
A comprehensive scan shows that a number of approximately 7×10^4 models in the class under consideration satisfy all criteria. Among them we have 9×10^3 super no-scale models. A tedious numerical calculation leads to 1792 models with $V_0 > 0$.



One loop potentials: Super no scale model potentials

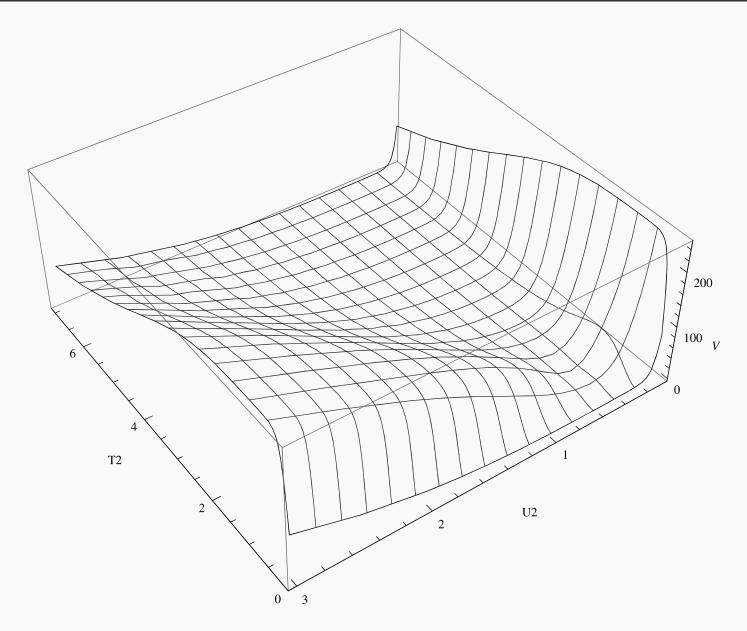


One loop potentials: Super no scale model potentials



One loop potential as a function of the T_2 , U_2 moduli.

One loop potentials: Super no scale models



One loop potential as a function of the T_2 , U_2 moduli.

Gauge coupling Running - Thresholds

The gauge coupling running is calculable in the context of string theory. It turns out that they depend on the compactification moduli. At the one loop level

$$\frac{16\pi^2}{g_i^2(\mu)} = k_a \frac{16\pi^2}{g_s^2} + b_a \log \frac{M_s^2}{\mu^2} + \Delta_a$$

where $M_S = g_S M_P$, $M_P = 1/\sqrt{32G_N}$.

 $b_a \leftrightarrow \text{Massless modes} \quad \Delta_a \leftrightarrow \text{Massive modes}$

$$\Delta_a = \Delta_a'(t_i) + \hat{\Delta}_a$$

see e.g. L. J. Dixon, V. Kaplunovsky and J. Louis (1991), I. Antoniadis, E. Gava, K.S. Narain (1992), C. Angelantonj, I. Florakis and M. Tsulaia (2014), Florakis (2015)

Decompactification problem

$$\Delta'_{a} - \Delta'_{b} = \sum_{i} \left\{ -\alpha^{i}_{ab} \log \left[T_{2}^{i} U_{2}^{i} | \eta(T^{i}) \eta(U^{i})|^{4} \right] -\beta^{i}_{ab} \log \left[T_{2}^{i} U_{2}^{i} | \vartheta_{4}(T^{i}) \vartheta_{2}(U^{i})|^{4} \right] -\gamma^{i}_{ab} \log \left[|\hat{j}_{2}(T^{i}/2) - \hat{j}_{2}(U^{i})|^{4} |j_{2}(U^{i}) - 24|^{4} \right] \right\},$$

 $\alpha^i_{ab}, \beta^i_{ab}, \gamma^i_{ab}$ model dependent coefficients. The dominant growth at $T^i_2\gg 1$

$$\Delta_a' = \alpha_a^i \left(\frac{\pi}{3} T_2^i - \log T_2^i \right) + \dots,$$

Solutions ? : $a_a^i = 0, ...$

Antoniadis (1990)

E. Kiritsis, C. Kounnas, P.M. Petropoulos, J. R. (1996)

Computation of the thresholds

The dominant moduli dependent contribution is

$$\Delta_a' = -\frac{k_a}{48} \, \mathsf{Y} + \hat{\beta}_a \, \Delta \,,$$

where the universal part Y is defined as

$$Y = \int_{\mathcal{F}} \frac{d^2 \tau}{\tau_2} \, \Gamma_{2,2}(T,U) \left(\frac{\hat{\bar{E}}_2 \bar{E}_4 \bar{E}_6 - \bar{E}_4^3}{\bar{\Delta}} + 1008 \right) \,,$$

$$\Delta = \int_{\mathcal{F}} \frac{d^2\tau}{\tau_2} \, \Gamma_{2,2}(T,U) = -\log\left[T_2 U_2 \, |\eta(T) \, \eta(U)|^4\right] \, .$$

At the limit $T_2 \gg 1$

$$Y = 48\pi T_2 + \mathcal{O}(T_2^{-1})$$
, $\Delta = \frac{\pi}{3}T_2 - \log T_2 + \mathcal{O}(e^{-2\pi T_2})$

and finally

$$\Delta_a = \left(\frac{\hat{\beta}_a}{3} - k_a\right) \pi T_2 + \mathcal{O}(\log T_2).$$

Computation of the thresholds

A comprehensive scan over a class of 7×10^4 models with $SO(10) \times SO(8)^2 \times U(1)^2$ gauge symmetry yields for the non-abelian gauge couplings

\hat{b}_{10}	\hat{b}_8	$\hat{b}_{8'}$	# of models	%
3	3	3	29456	38.5
9	-3	-3	15840	20.7
-3	9	9	14000	18.3
•	•	•	•••	22.5

In a big class of vacua we have $\hat{\beta}_a = 3k_a$ (decompactification condition), hence there is no decompactification problem for the gauge couplings g_{10}, g_8, g_8'

$$\Delta_a = \underbrace{\hat{\beta}_a}_{3} \times R_a \pi T_2 + \mathcal{O}(\log T_2)$$

Gauge coupling running

For models satisfying the decompactification condition $\hat{\beta}_a = 3k_a$ the coupling running is

$$\frac{16\pi^2}{g_a^2(\mu)} = k_a \frac{16\pi^2}{g_s^2} + \beta_a \log \frac{M_s^2}{\mu^2} + \beta_a' \log \left(\frac{2e^{1-\gamma}}{3\pi\sqrt{3}} \frac{M_{KK}^2}{M_s^2}\right) + \dots$$

Here, γ is the Euler-Mascheroni constant, $M_{KK}=1/\sqrt{T_2}$ is the Kaluza-Klein scale. $\beta_a=b_a^{(1)}+b_a^{(2)}+b_a^{(3)}$ and $\beta_a'=b_a^{(1)}+b_a^{(2)}$ with $b_a^{(1)}=\hat{\beta}_a$

A Standard Model scenario

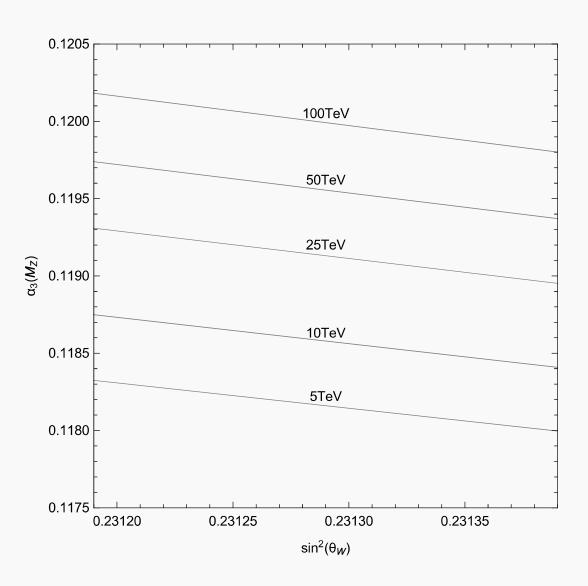
$$\frac{k_2 + k_Y}{\alpha_S} = \frac{1}{\alpha_{em}} - \frac{\beta_2 + \beta_Y}{4\pi} \log \frac{M_S^2}{M_Z^2} - \frac{\beta_2' + \beta_Y'}{4\pi} \log \left(\frac{2e^{1-\gamma}}{3\pi\sqrt{3}} \frac{M_{KK}^2}{M_S^2} \right)$$

$$\sin^{2}\theta_{W} = \frac{k_{2}}{k_{2} + k_{Y}} + \frac{\alpha_{em}}{4\pi} \left[\frac{k_{Y}\beta_{2} - k_{2}\beta_{Y}}{k_{2} + kY} \log \frac{M_{s}^{2}}{M_{Z}^{2}} + \frac{k_{Y}\beta_{2}' - k_{2}\beta_{Y}'}{k_{2} + kY} \log \left(\frac{2e^{1-\gamma}}{3\pi\sqrt{3}} \frac{M_{KK}^{2}}{M_{s}^{2}} \right) \right]$$

$$\frac{1}{\alpha_3(M_Z)} = \frac{k_3}{\alpha_{\text{em}}(k_2 + k_Y)} + \frac{1}{4\pi} \left[\left(\beta_3 - \frac{k_3(\beta_2 + \beta_Y)}{k_2 + k_Y} \right) \log \frac{M_S^2}{M_Z^2} + \left(\beta_3' - \frac{k_3(\beta_2' + \beta_Y')}{k_2 + k_Y} \right) \log \left(\frac{2e^{1-\gamma}}{3\pi\sqrt{3}} \frac{M_{\text{KK}}^2}{M_S^2} \right) \right]$$

A Standard Model scenario

For
$$(\beta_Y, \beta_2, \beta_3) = (-7, -\frac{19}{6}, \frac{41}{6})$$
, $(k_Y, k_2, k_3) = (\frac{5}{3}, 1, 1)$ and $(\beta_Y', \beta_2', \beta_3') = (-\frac{15}{2}, -\frac{43}{6}, -\frac{23}{3})$.



Conclusions

We have analysed a class of non supersymmetric heterotic vacua where SUSY is spontaneously broken via the Scherk–Schwartz mechanism. In this context we have constructed semi-realistic models with the following interesting characteristics

- Fermion chirality
- Dynamical determination of supersymmetry breaking scale $M_{susy} \ll M_{Planck}$
- Exponentially small cosmological constant
- Finite gauge coupling running (no decompactification problem)
- These developments pave the way for non-supersymmetric string phenomenology (consider more realistic models e.g. Pati-Salam)